

# The ultra-steep diffuse emission observed in the cool-core cluster **RXJ 1720.1+2638** at low frequencies

**Nadia Biava**

Università di Bologna – Italy

Collaborators: Francesco de Gasperin, Annalisa Bonafede et al.

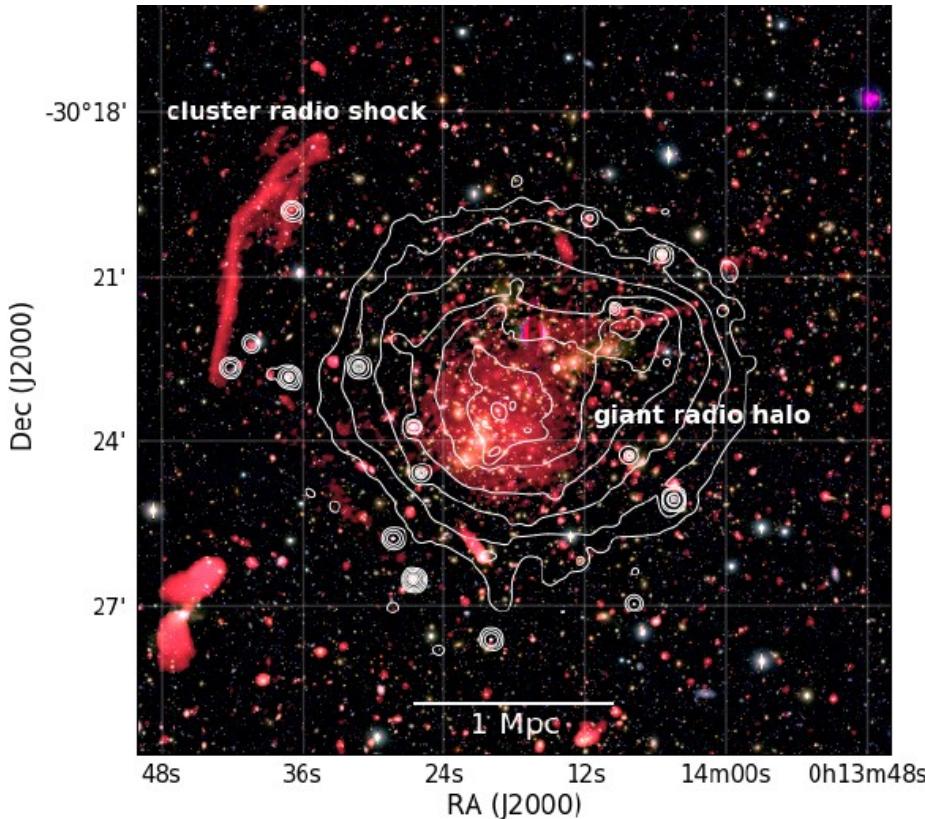


9 March 2021 – The LOFAR LBA mini-symposium

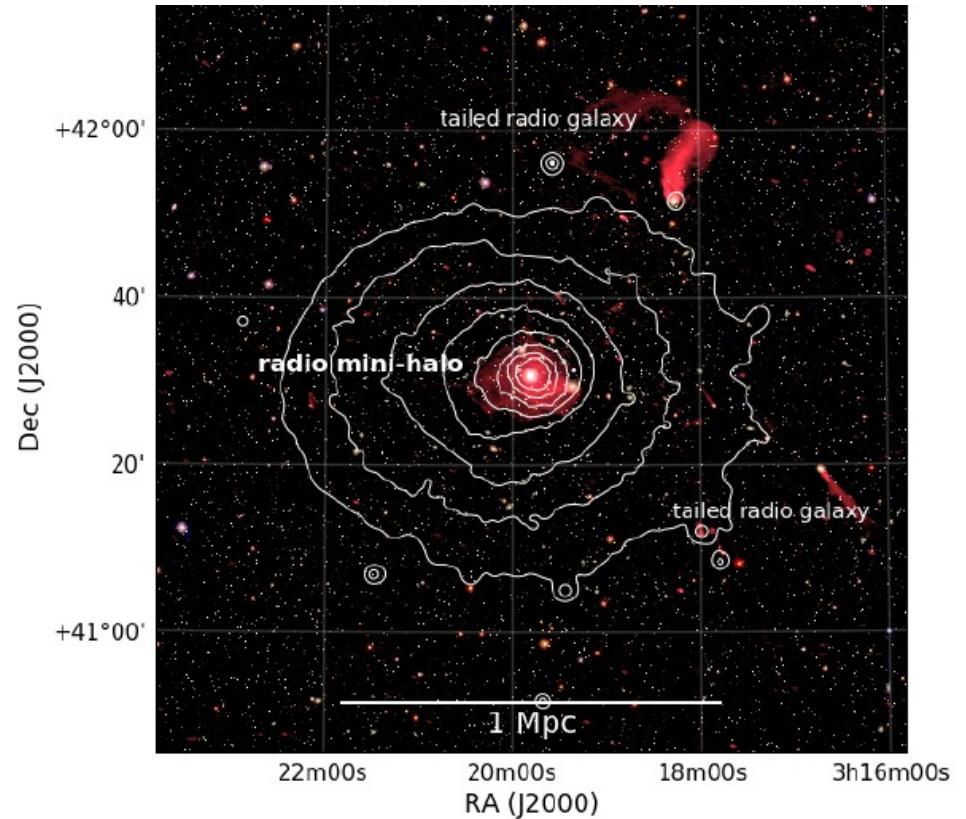


# Diffuse emission in galaxy clusters

Giant Halo



Mini halo



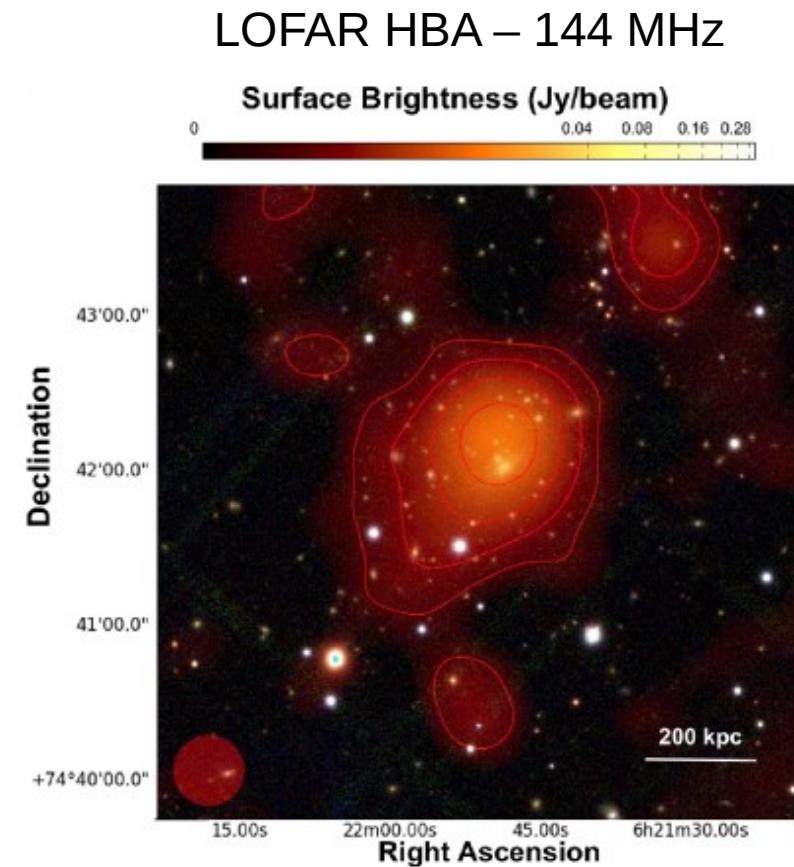
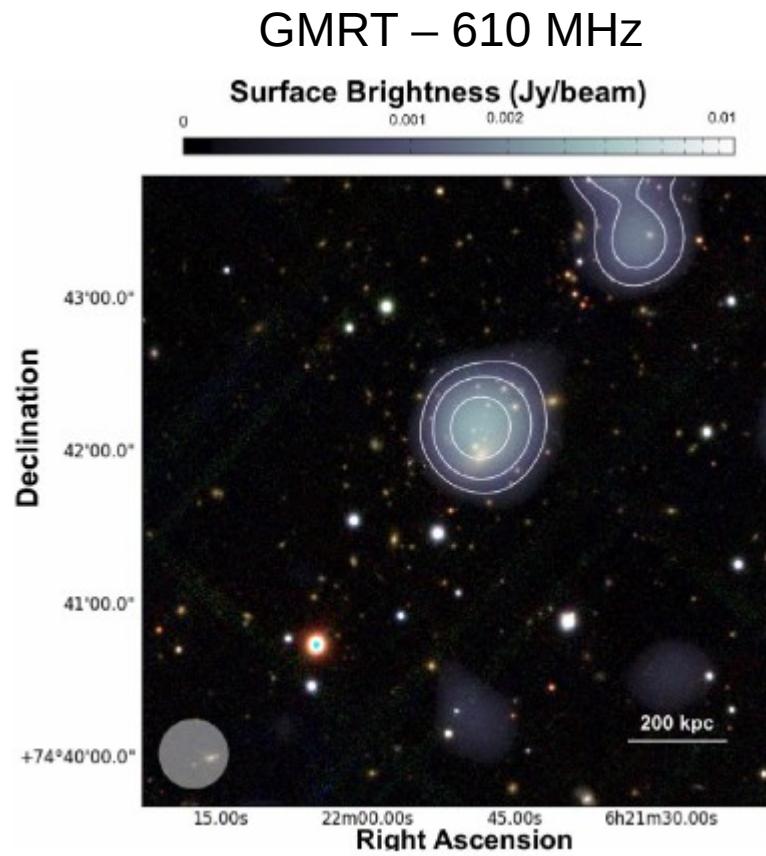
- Mpc size
- Merging clusters

- 100-500 kpc size
- Cool-core clusters

**Characterized by a steep radio spectrum ( $\alpha > 1$ )**

# A more complex picture..

- PSZ1G139.61+24



- Diffuse emission beyond the cluster core
- Steep spectrum  $\alpha > 1.5$

(Savini et al. 2018)

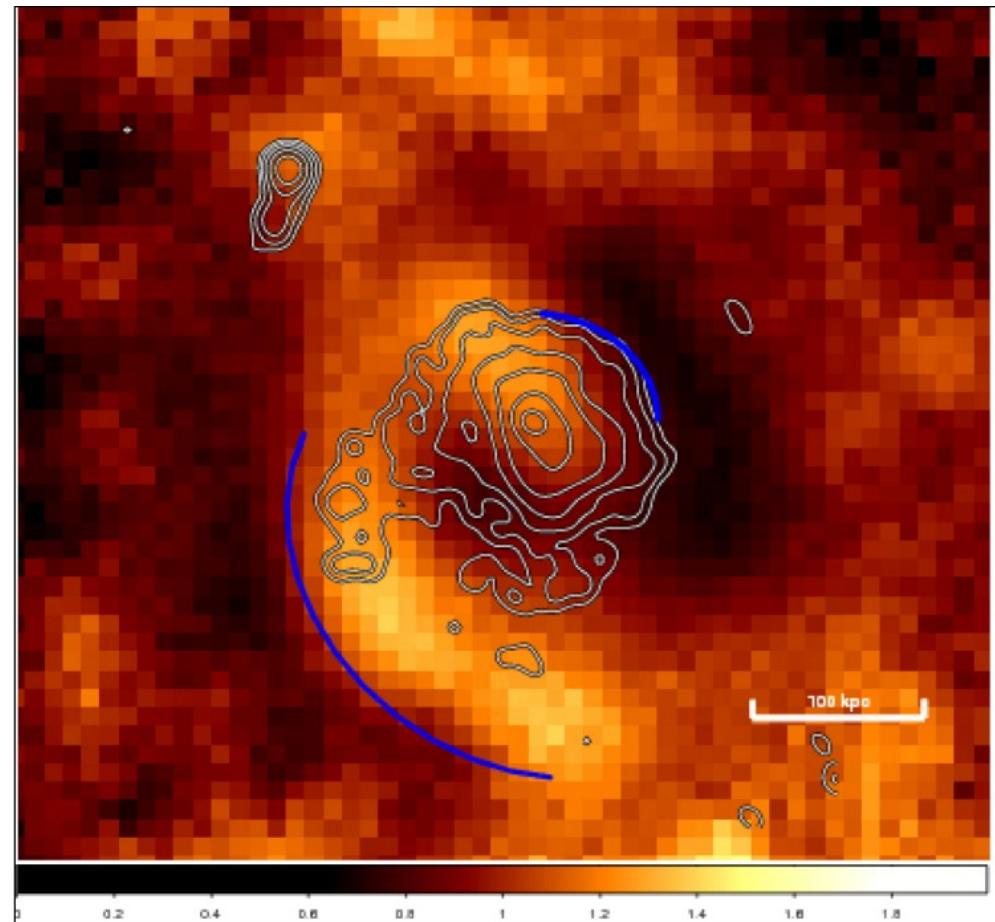
# RXJ 1720.1+2638

Cool-core  
galaxy cluster

$z = 0.16$

Mini halo + Cold fronts

Spatial correlation  
due to gas sloshing?

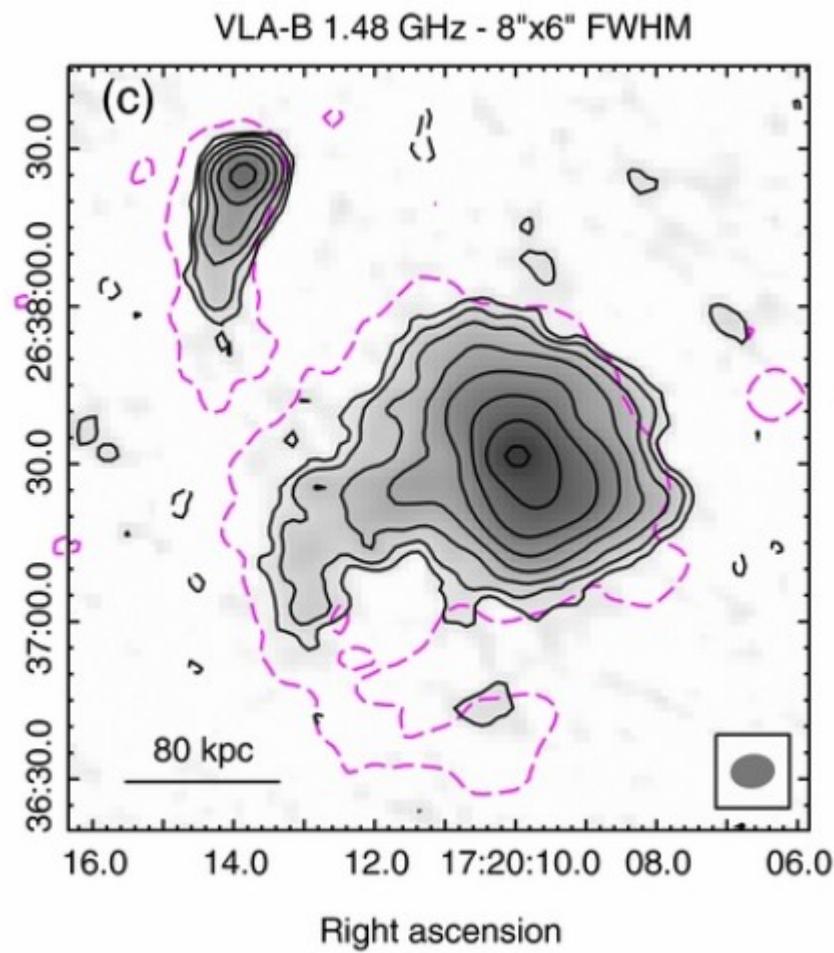


Ratio Chandra X-ray image in the [0.5, 2.5] keV band  
overlaid on the 1.5 GHz VLA-B contours.

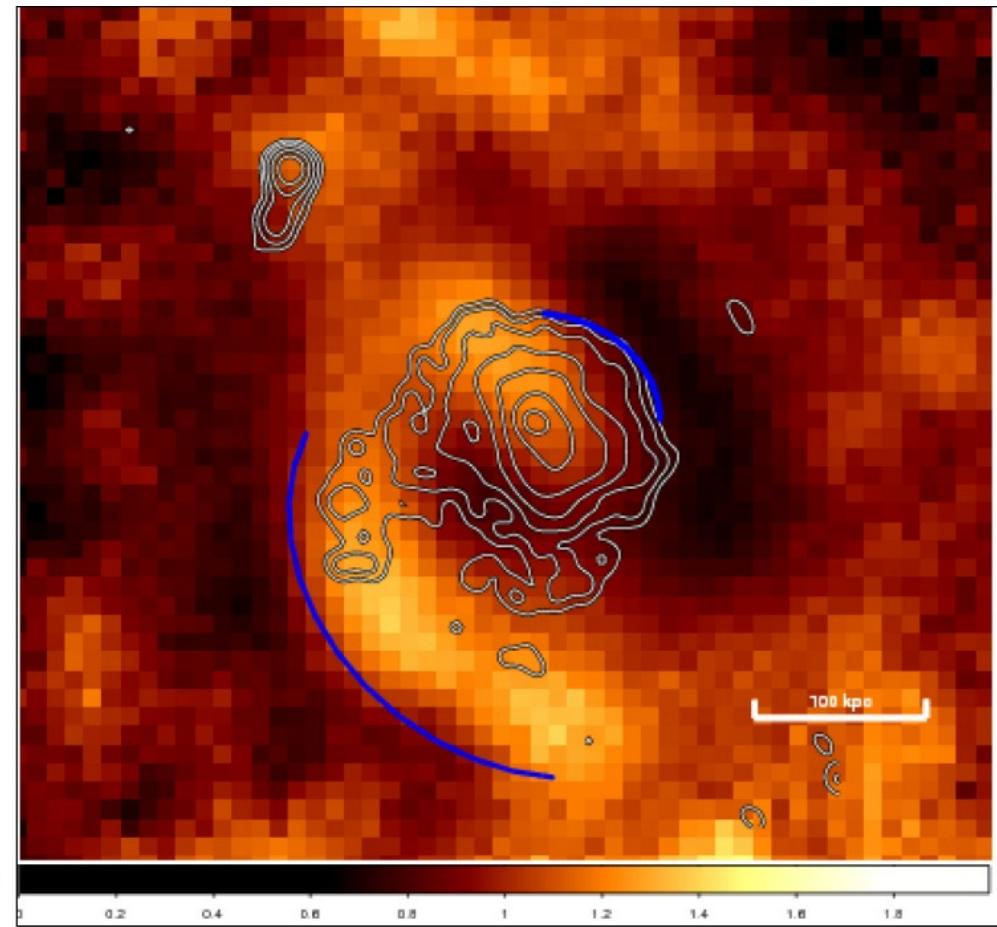
(Mazzotta & Giacintucci 2008)

# RXJ 1720.1+2638

Declination

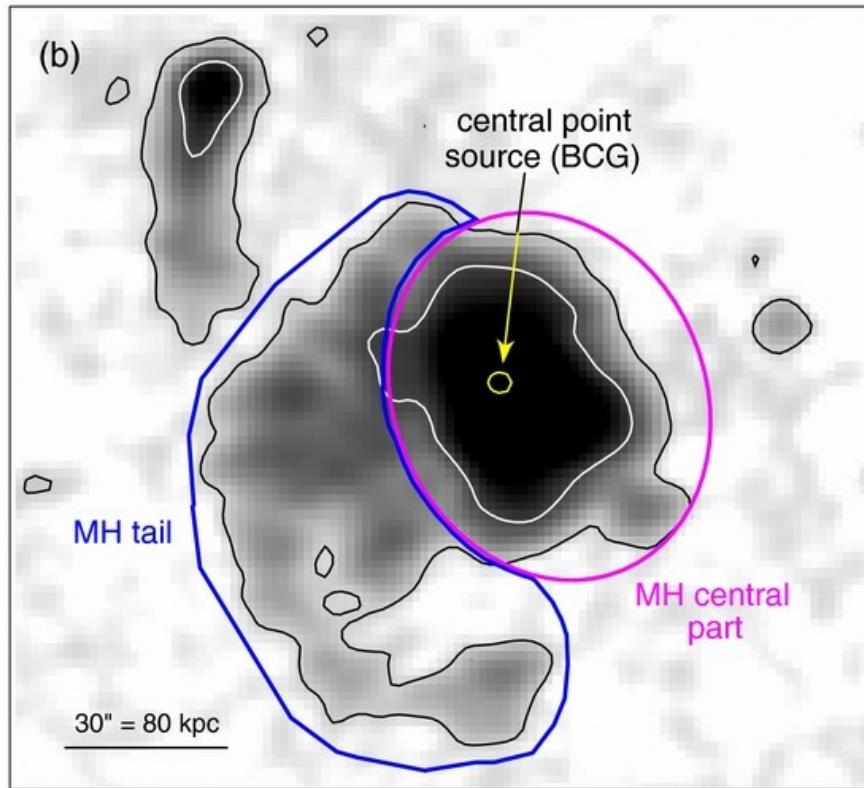


(Giacintucci et al. 2014)



(Mazzotta & Giacintucci 2008)

# RXJ 1720.1+2638



GMRT + VLA  
317 MHz – 4.86 GHz

Spectral index:

- MH central part

$$\alpha = 1.0 \pm 0.1$$

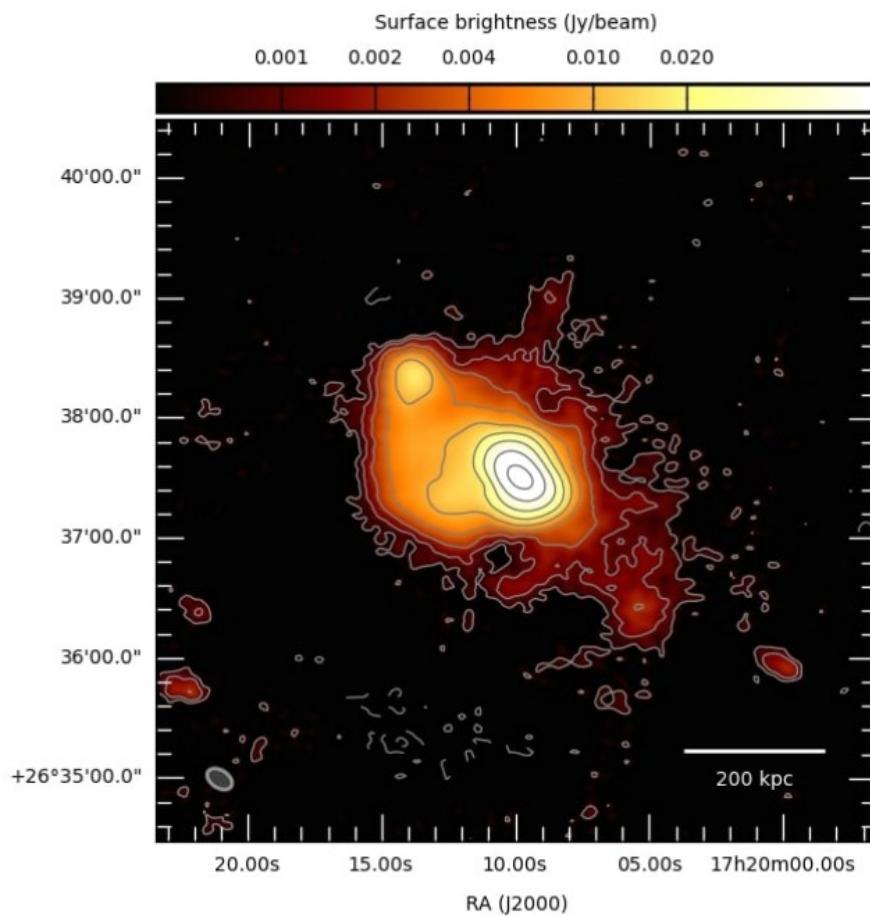
- MH tail

$$\alpha = 1.4 \pm 0.1$$

(Giacintucci et al. 2014)

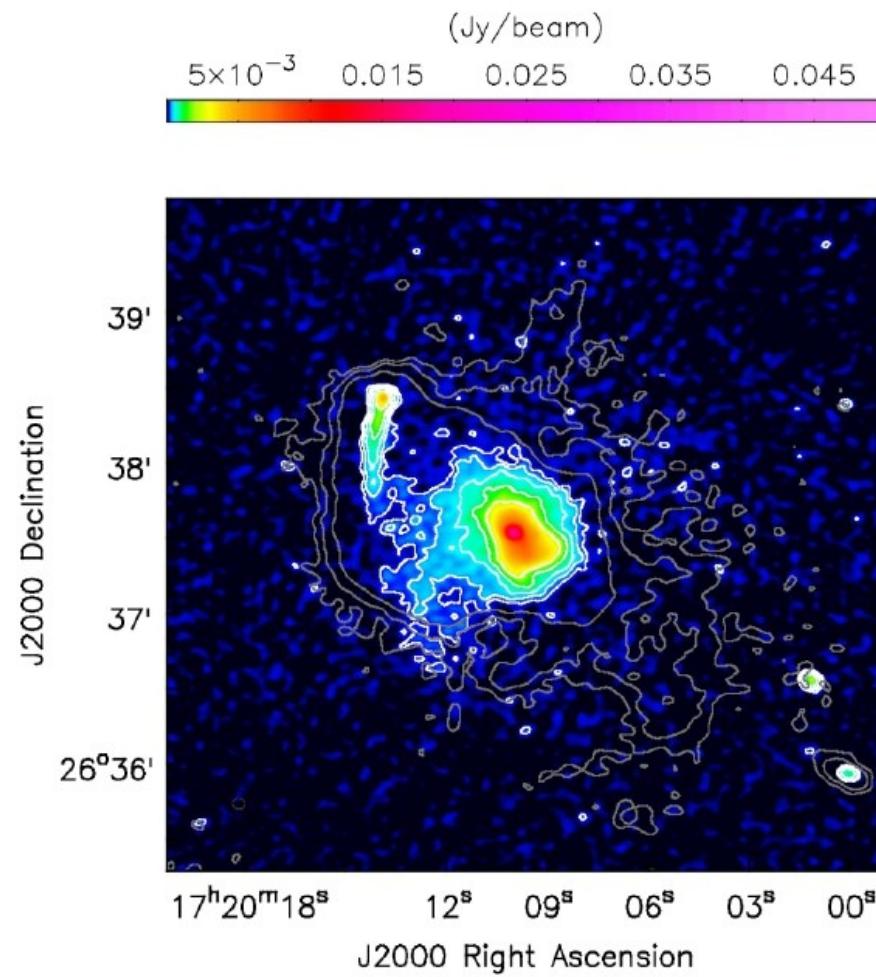
# RXJ 1720.1+2638 – LOFAR HBA

LOFAR HBA – 144 MHz



(Savini et al. 2019)

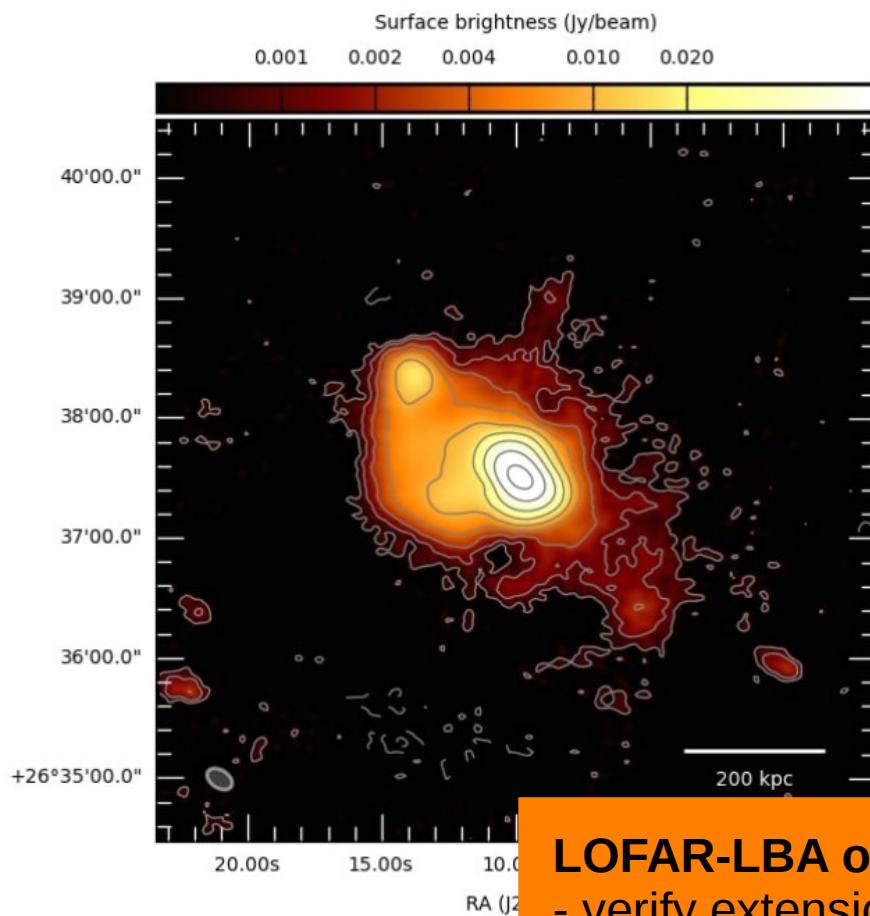
GMRT – 610 MHz



Diffuse emission beyond the cluster core, with  $\alpha > 2.1$

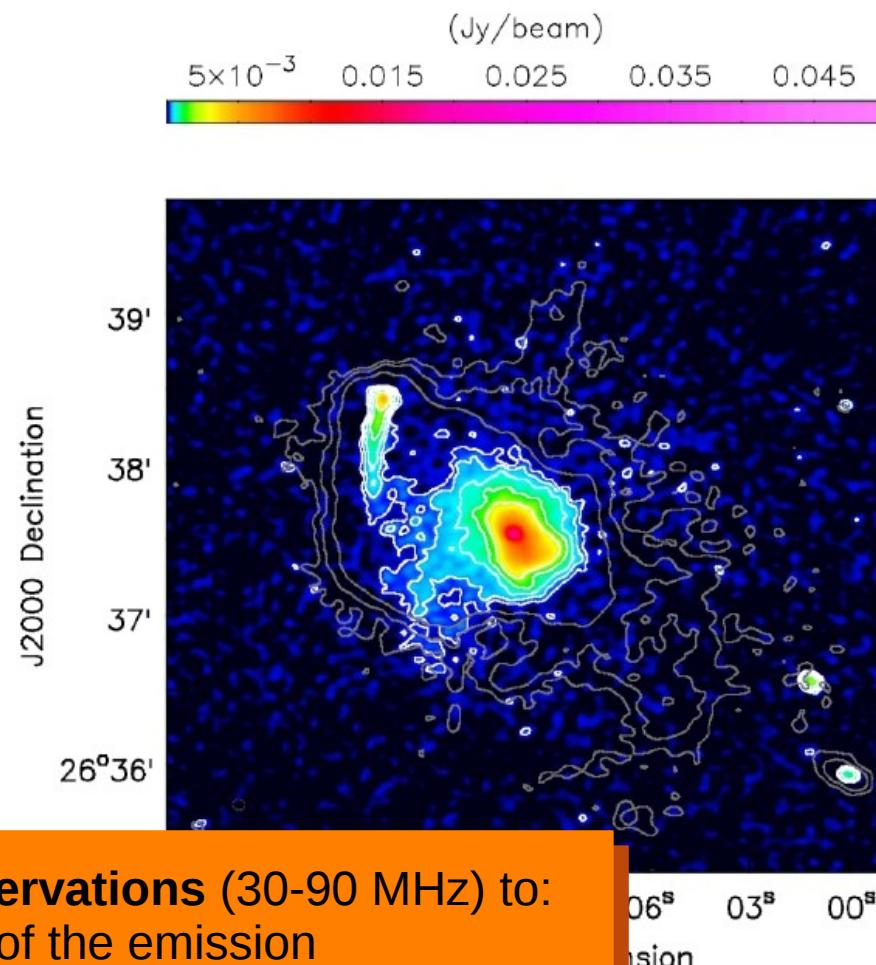
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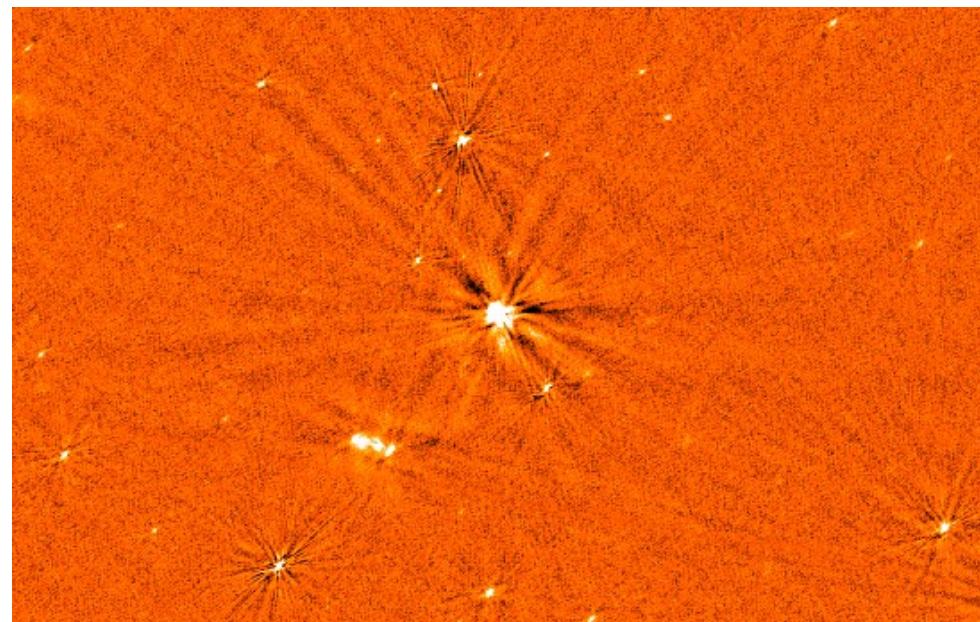
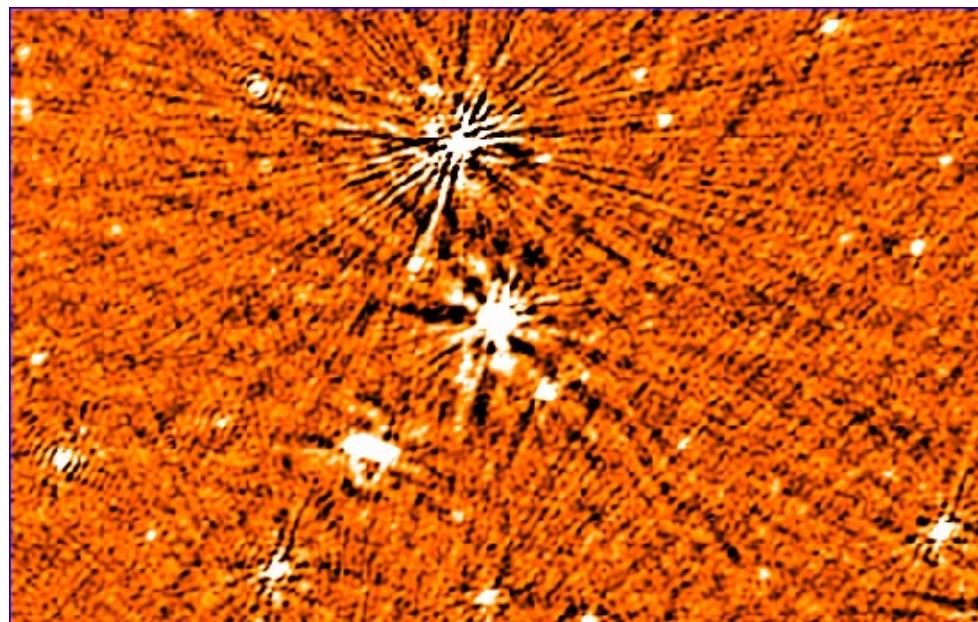
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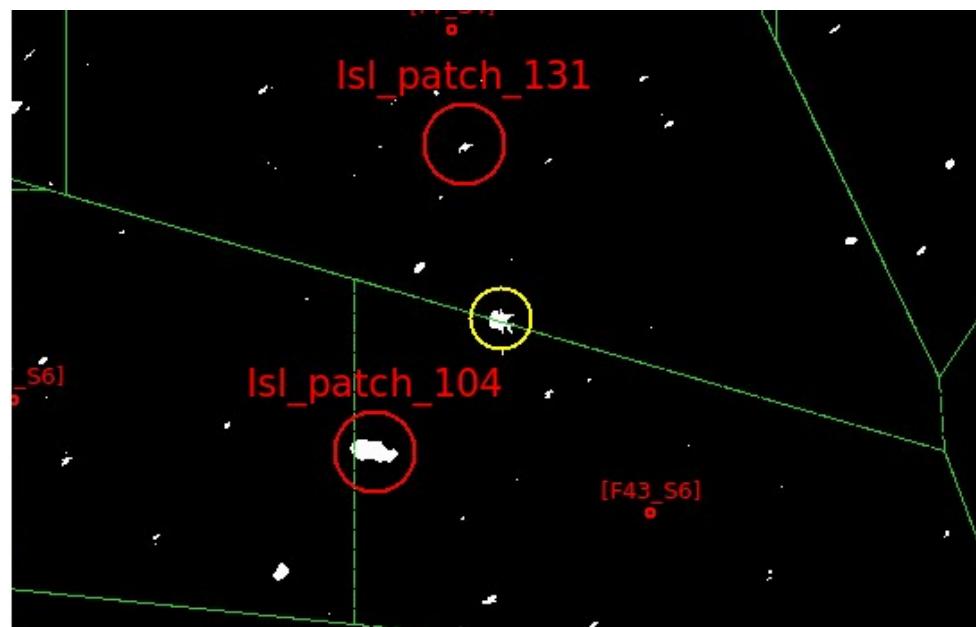
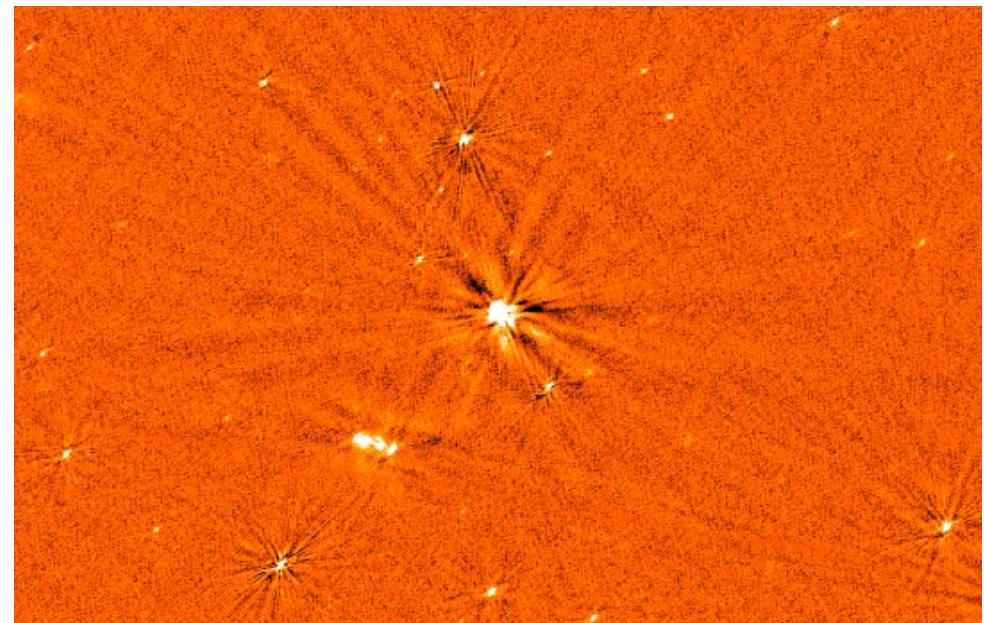
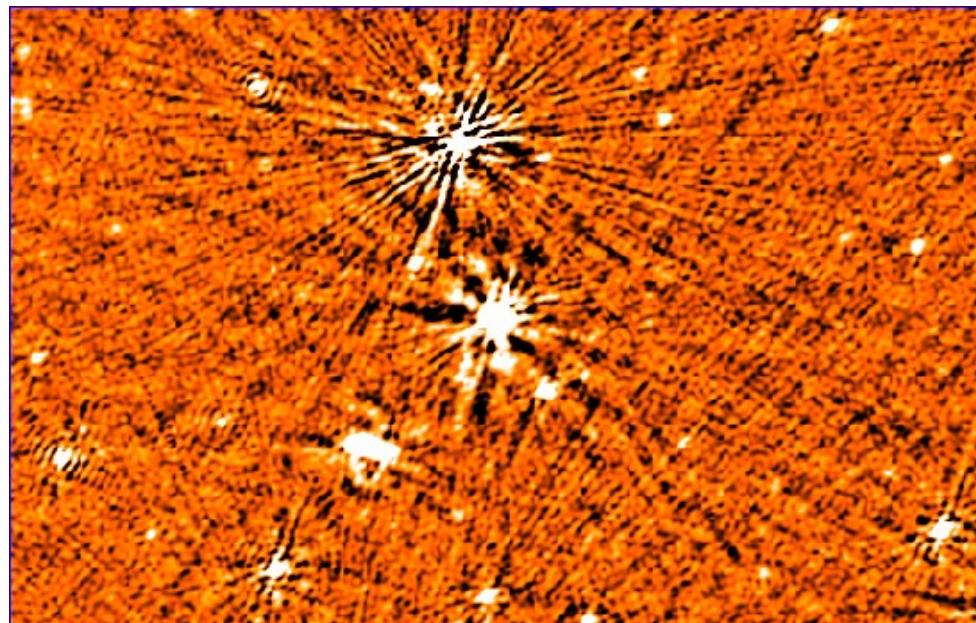
**LOFAR-LBA observations** (30-90 MHz) to:

- verify extension of the emission
  - create spectral index map of the whole halo
  - understand the origin of the emission
- 10 h awarded in cycle XX

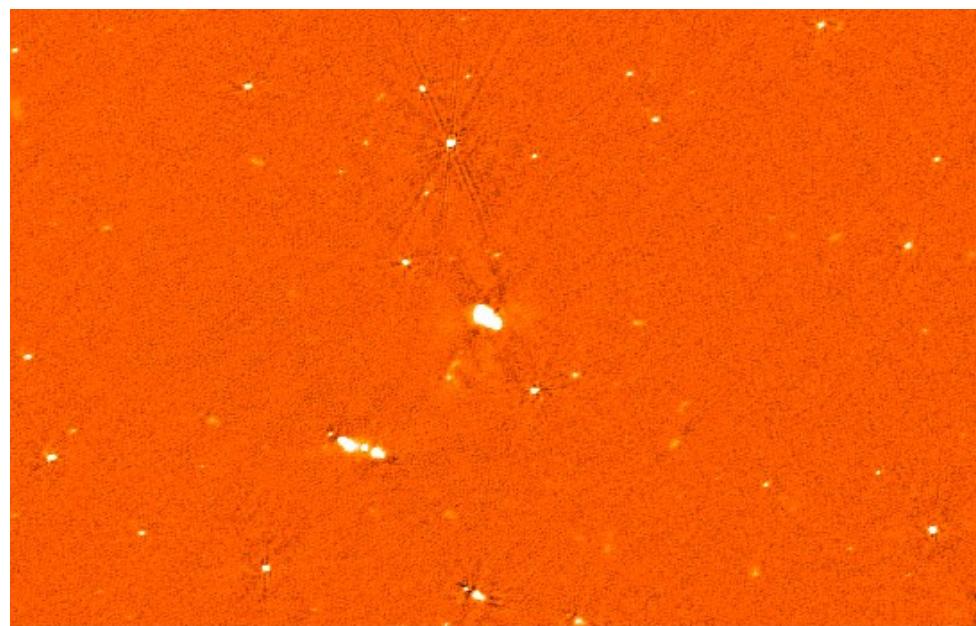
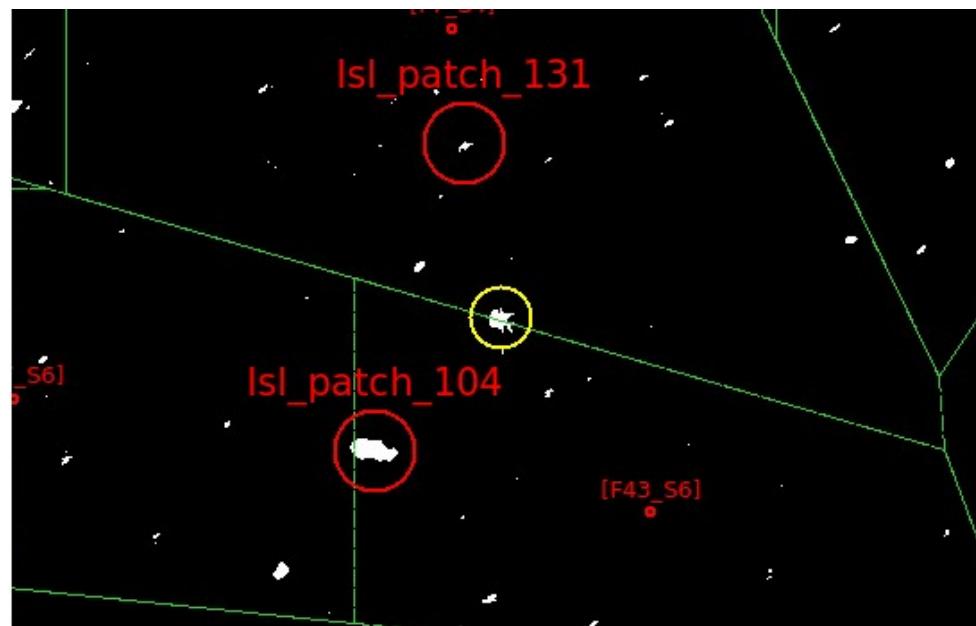
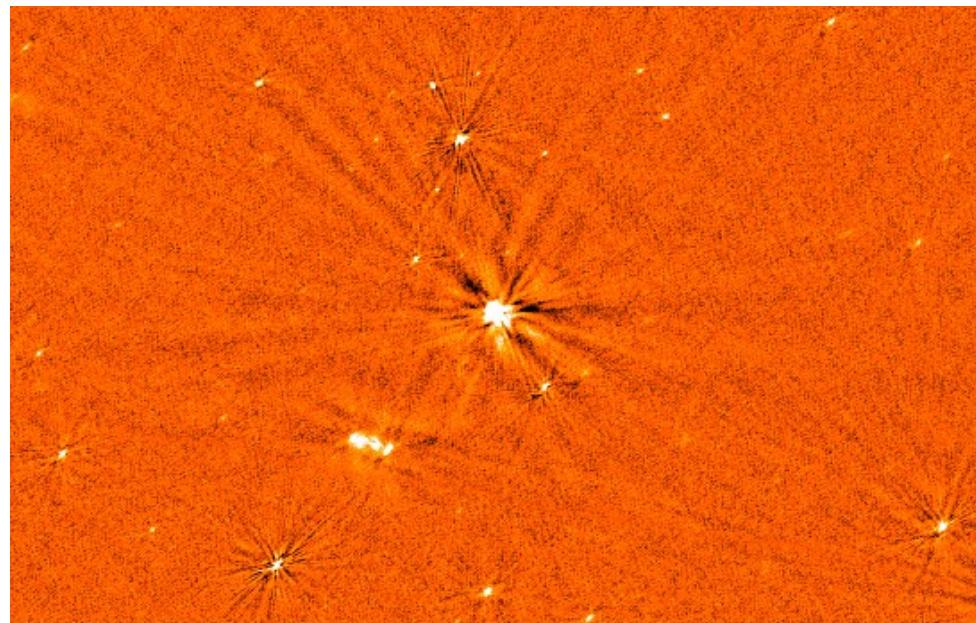
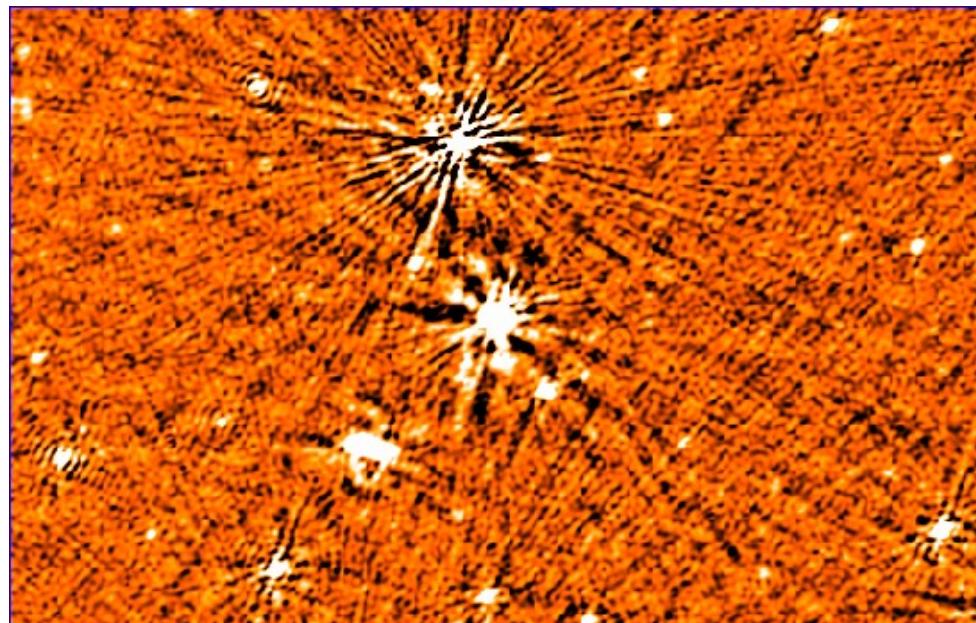
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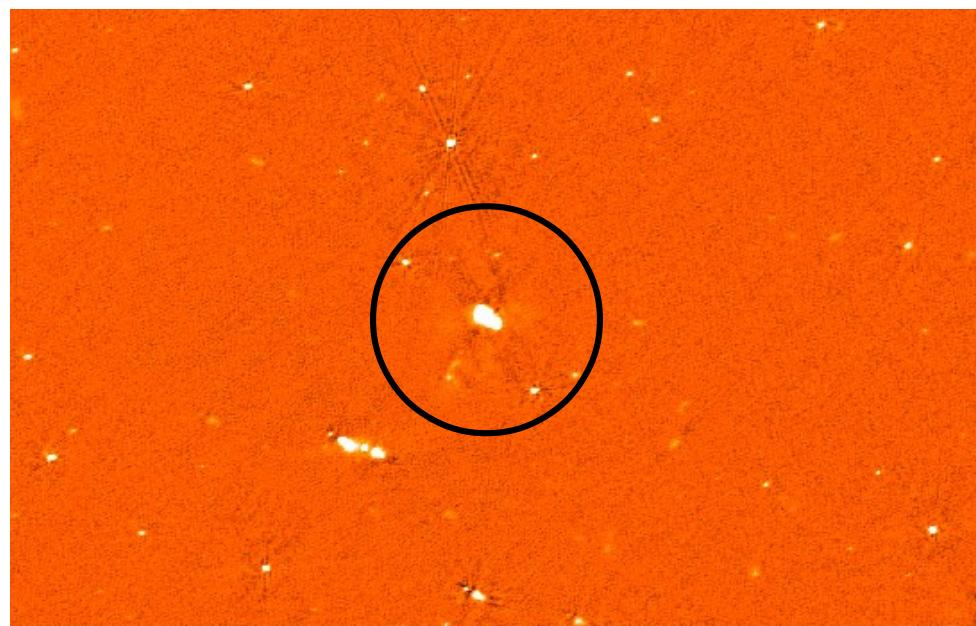
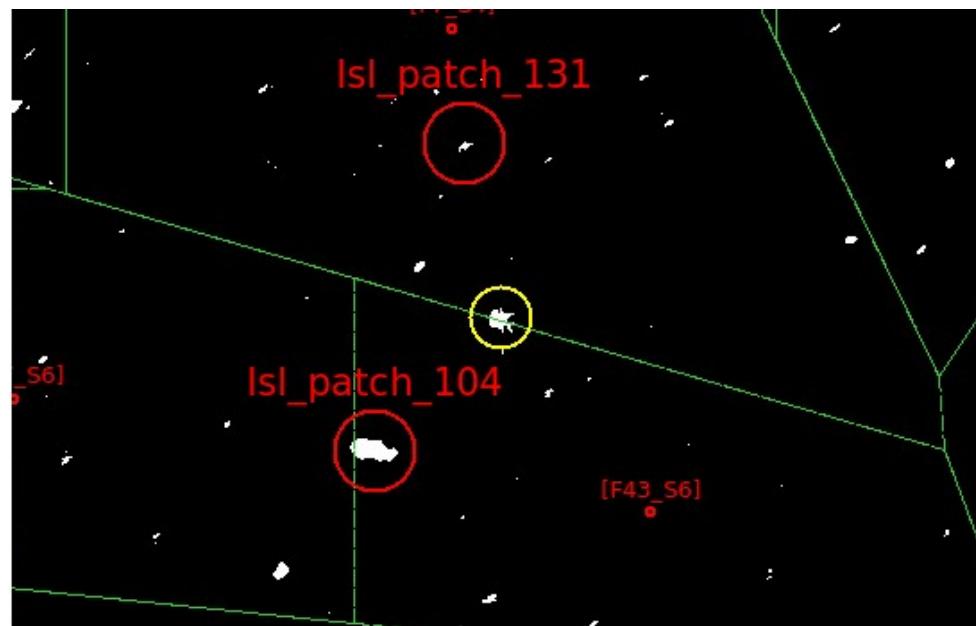
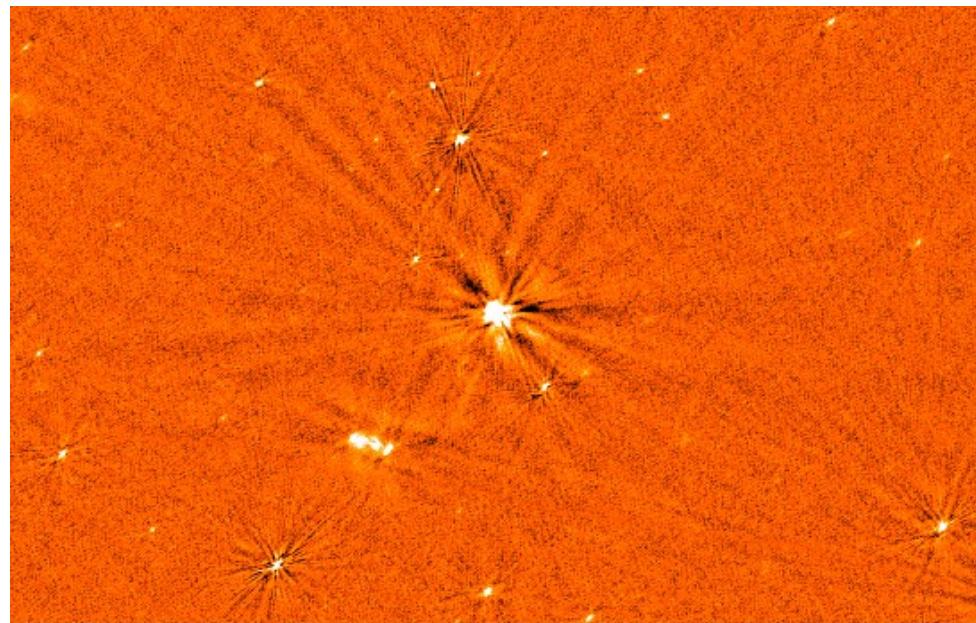
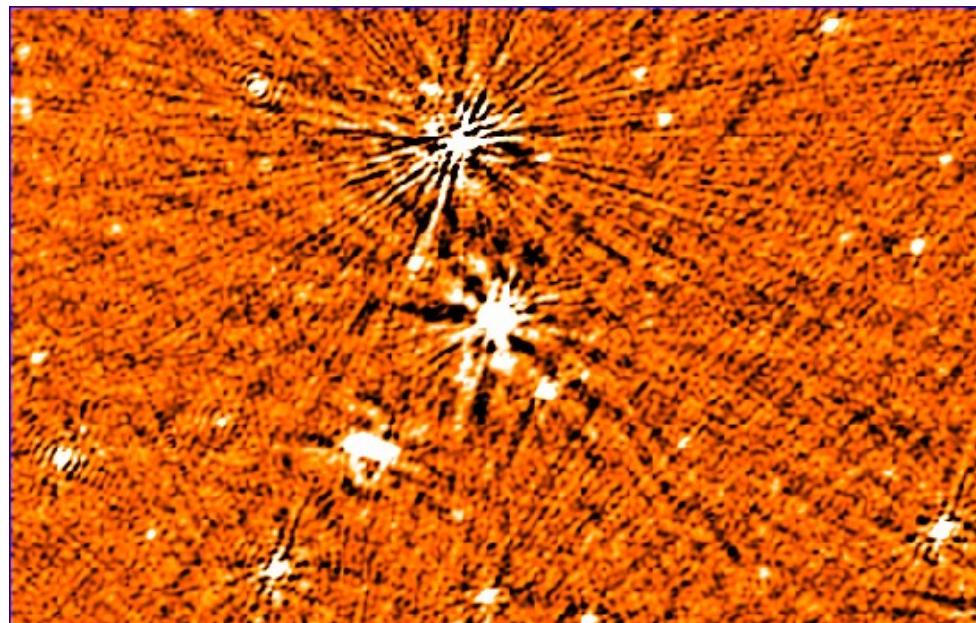
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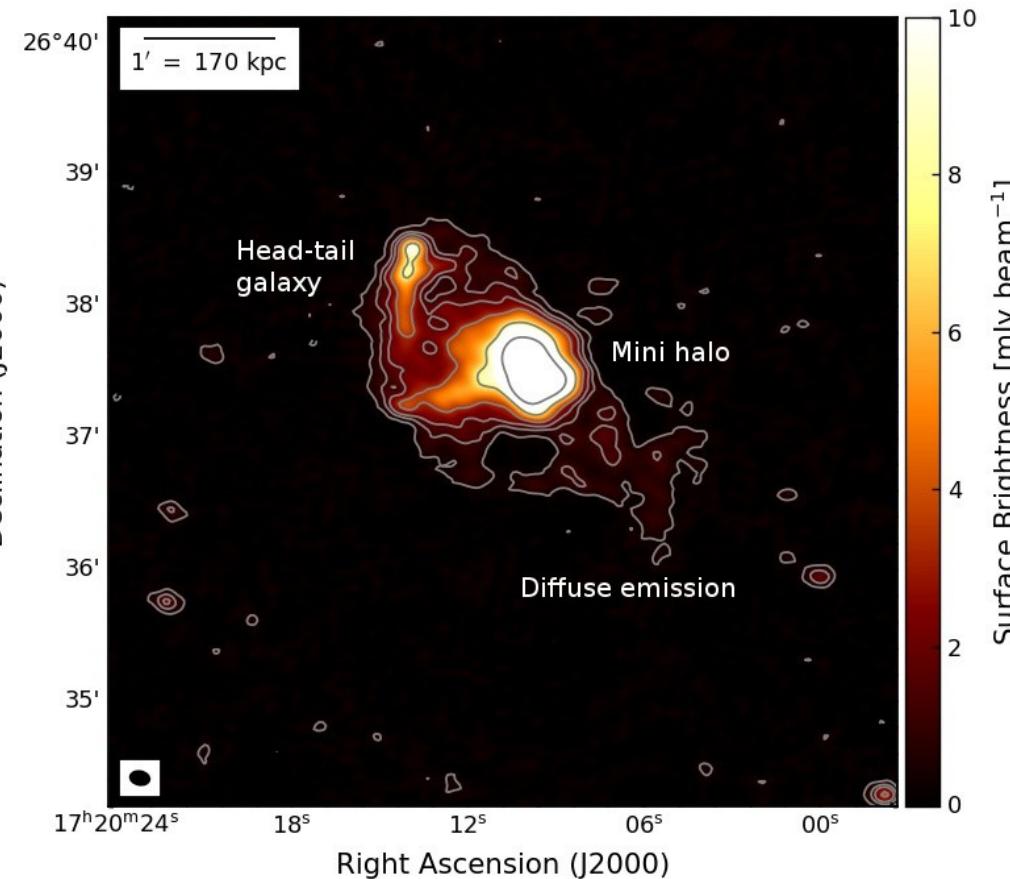


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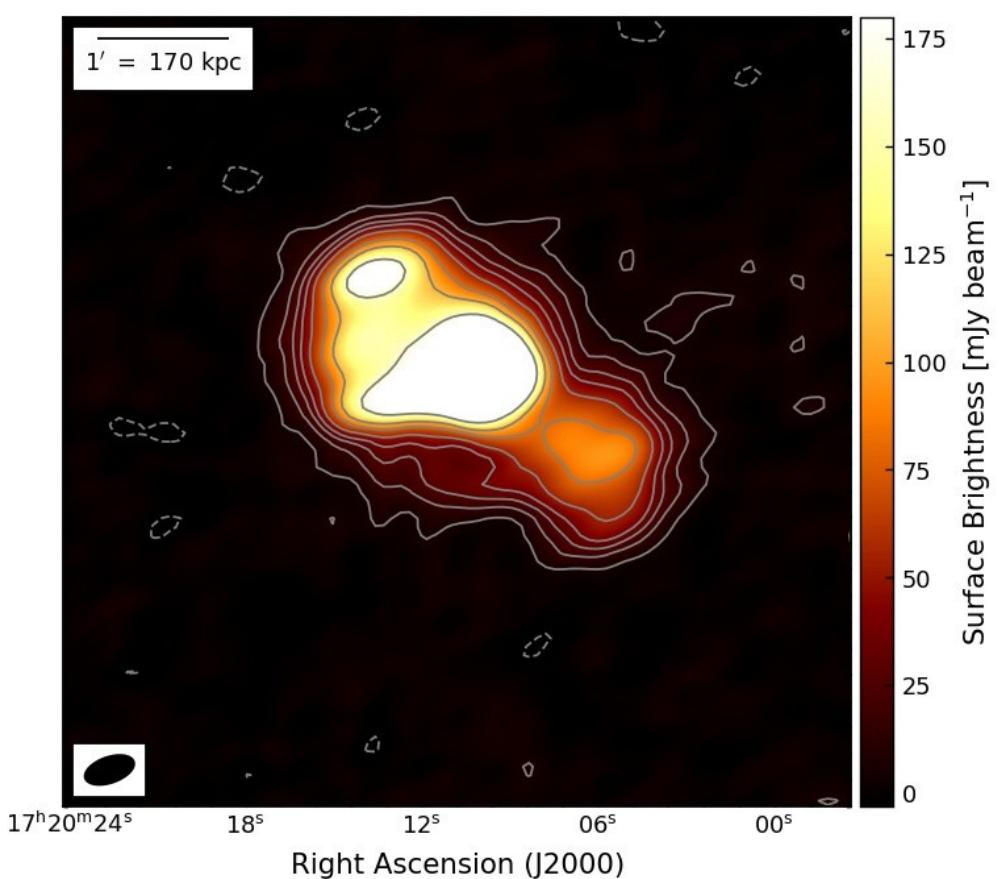
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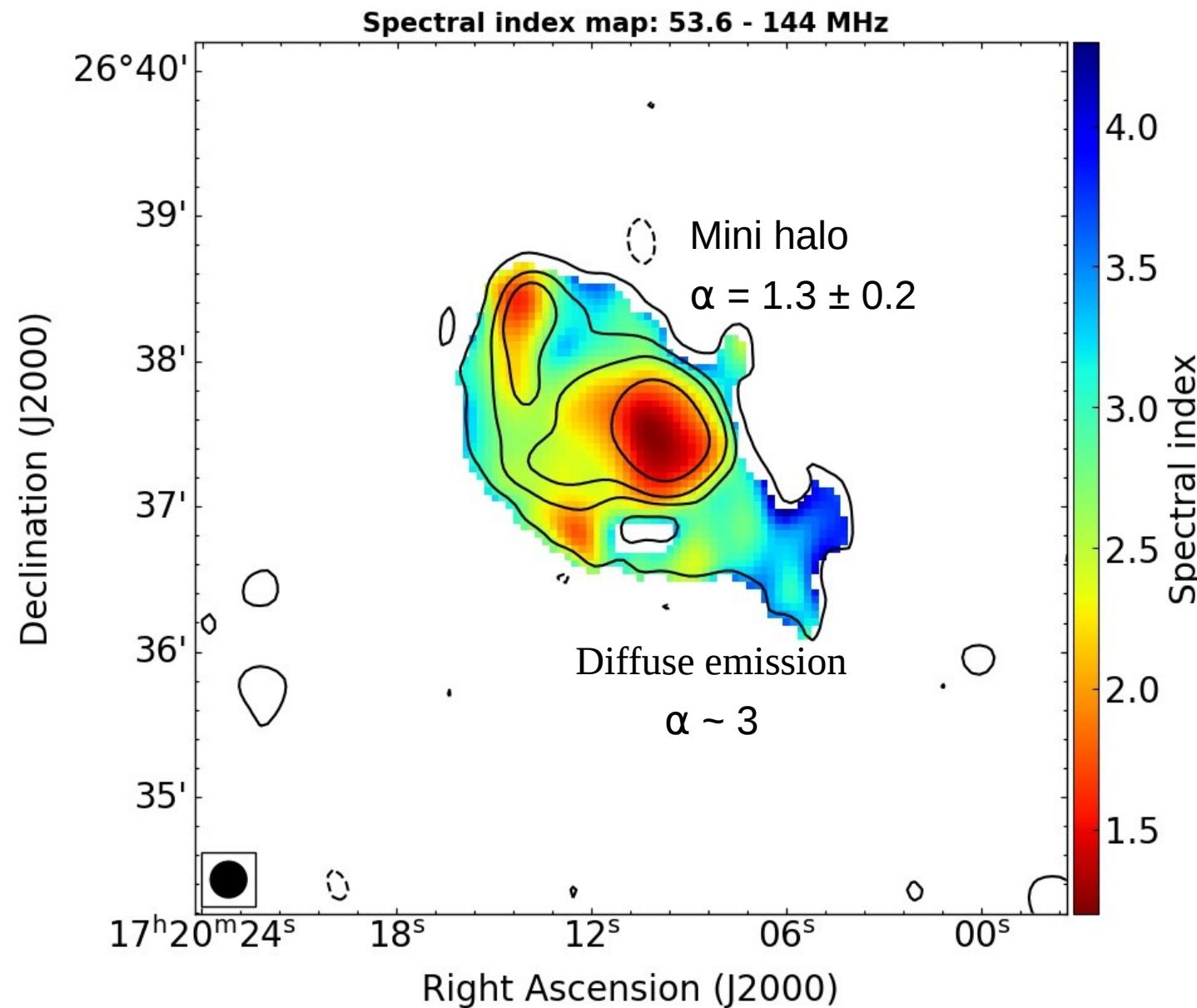
Res: 9" x 7"  
Rms:  $1.3 \times 10^{-1}$  mJy

LOFAR LBA – 53.6 MHz

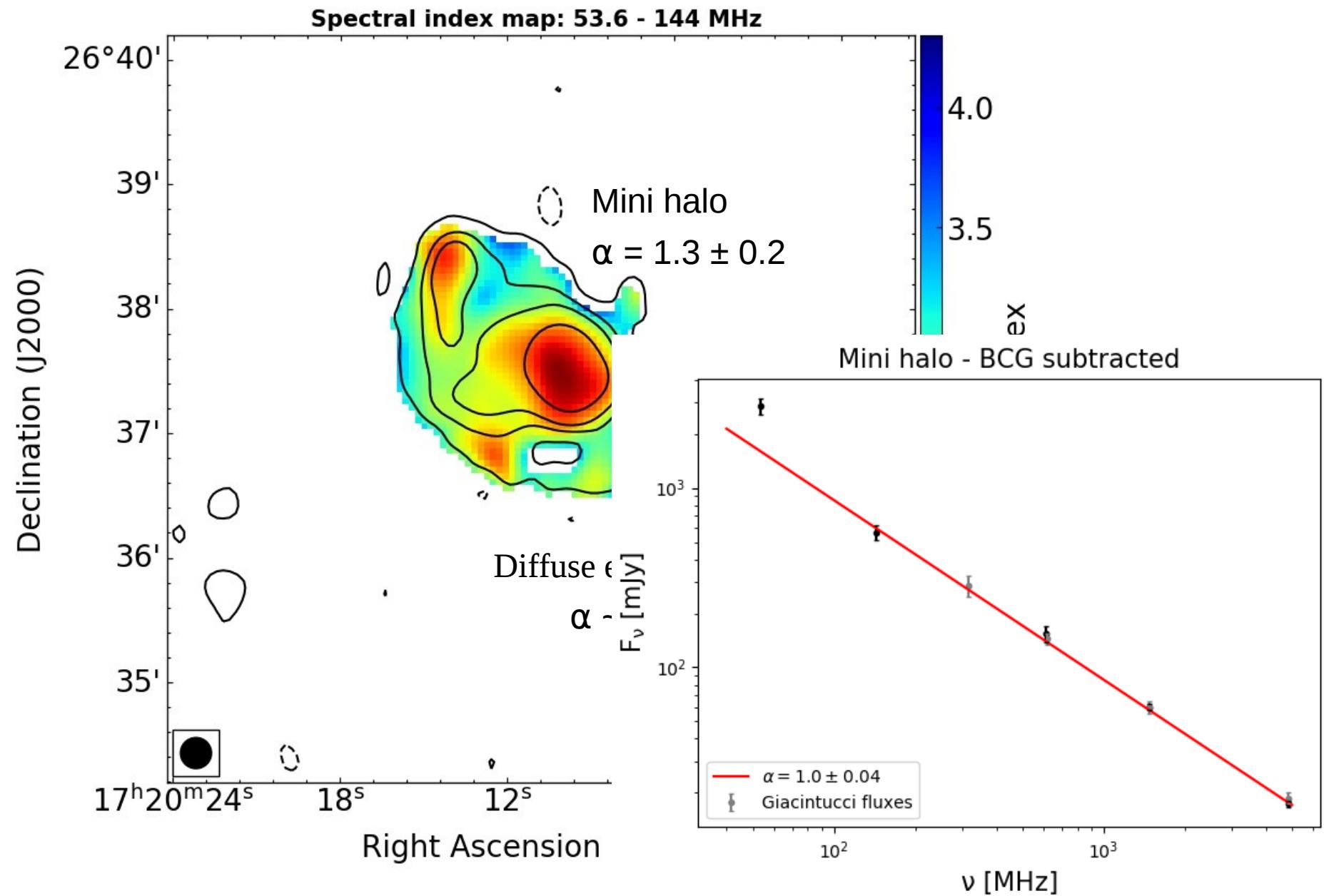


Res: 23" x 12"  
Rms: 1.8 mJy

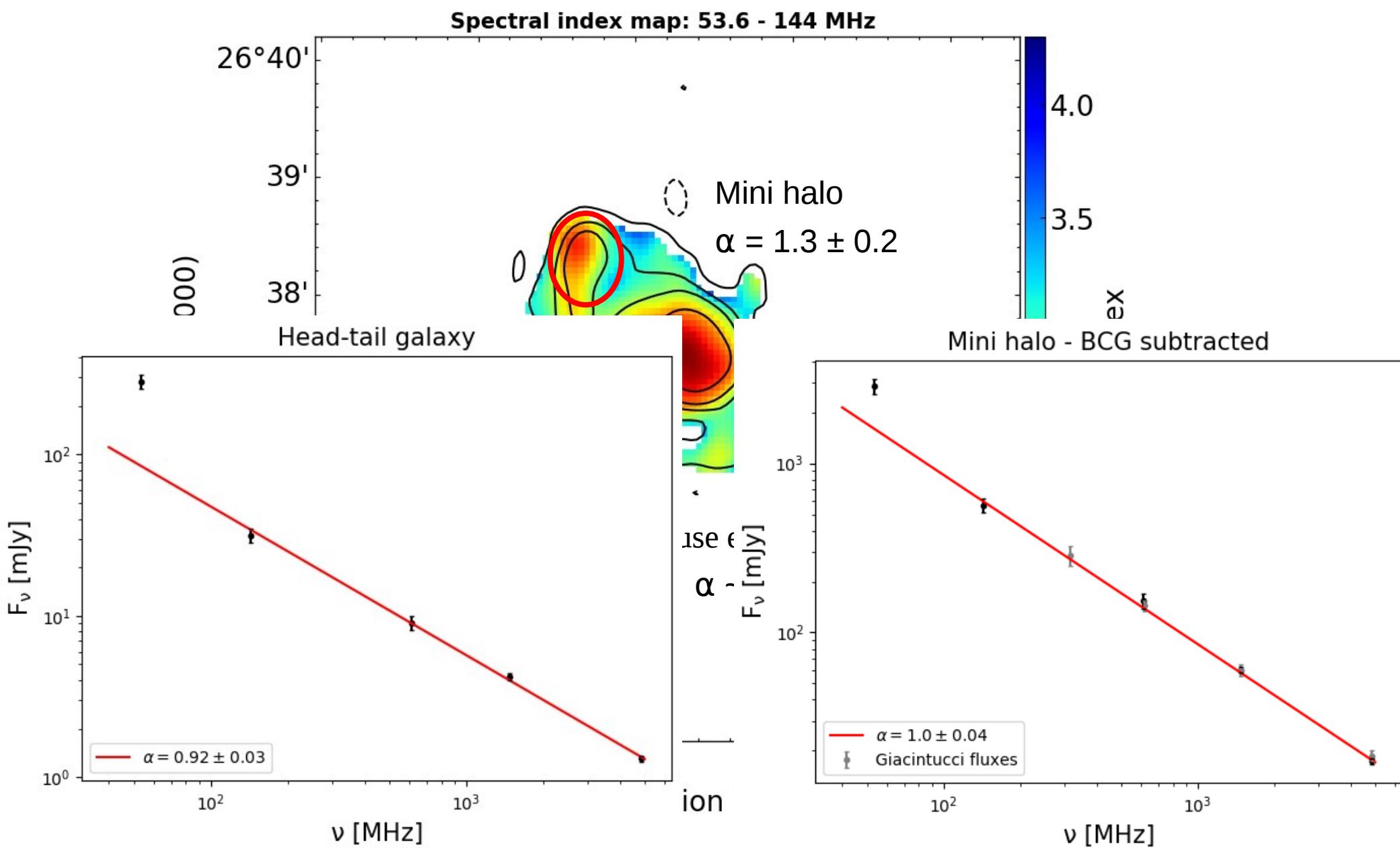
# Spectral index map



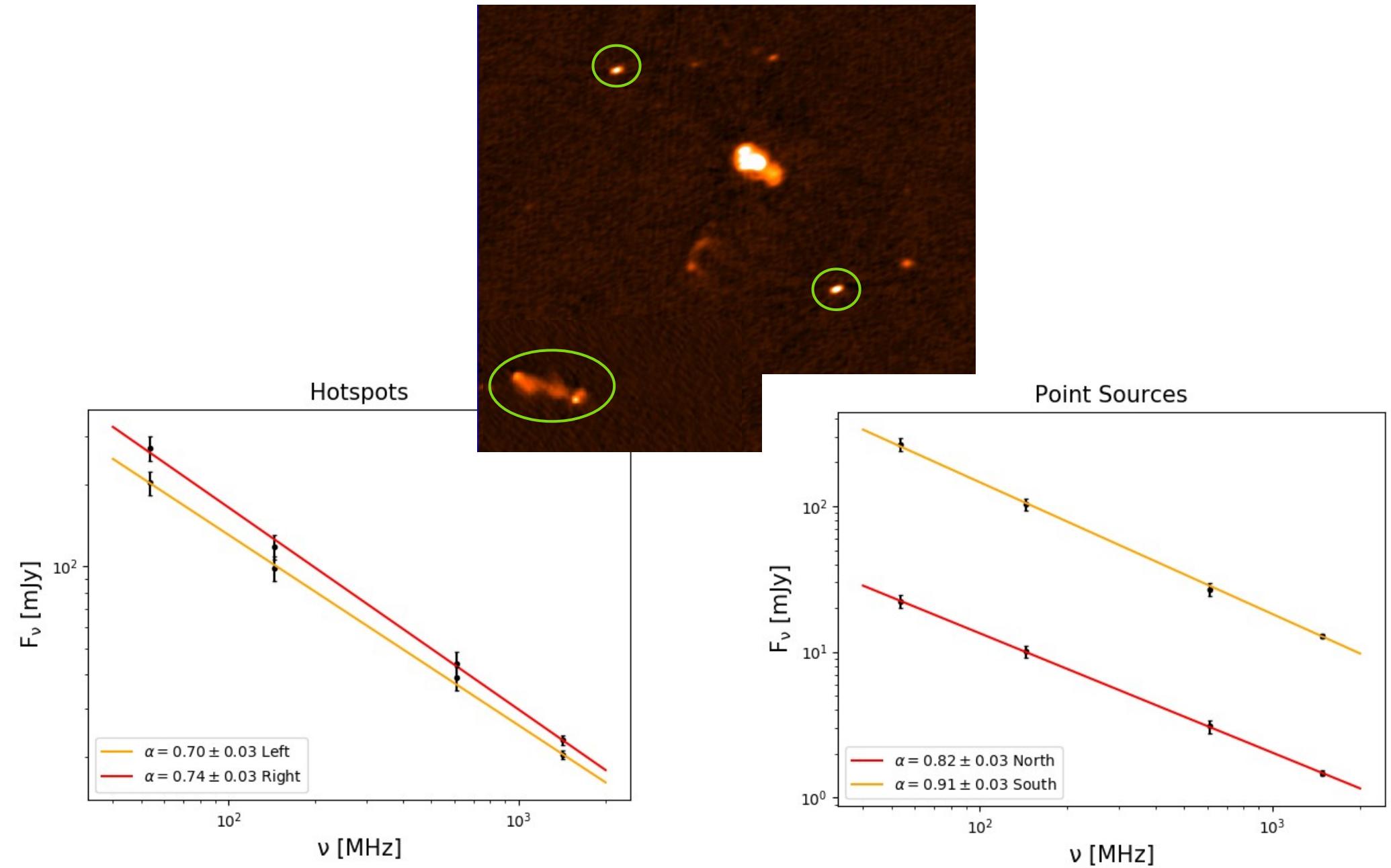
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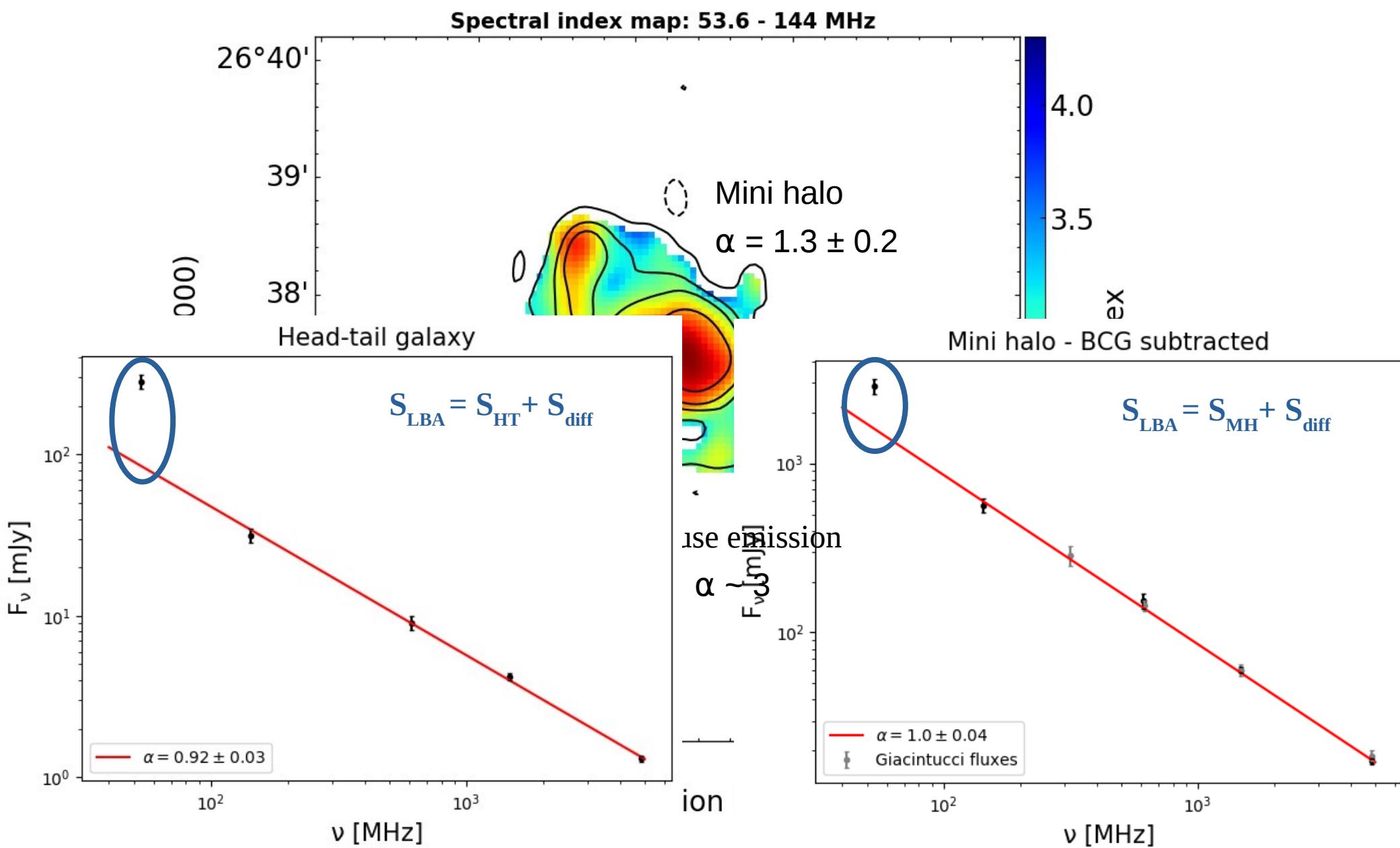
# Spectral index map



# Spectral index map



# Spectral index map



# Conclusions

- The diffuse emission detected with LOFAR has an ultra-steep spectrum.
- There is a net difference in the spectral index of the two components.

## Next steps

Analysis of X-ray data to check for a possible:

- spatial correlation between radio and X-ray surface brightness of the two components.
- cavity in correspondence of diffuse emission.

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**Thank you for the attention**