



Spectral modeling of 4C 43.15

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Observatory

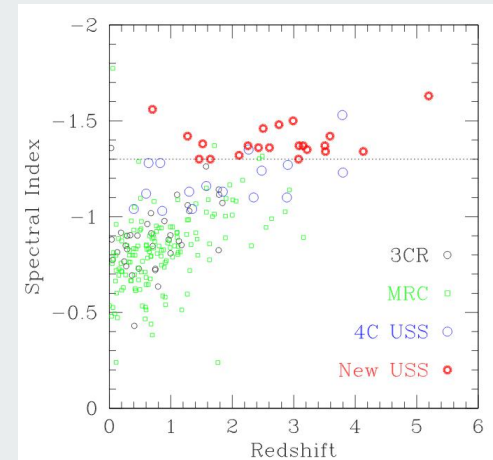


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High-z radio galaxies and the α -z correlation

- Important probes for galaxy evolution: e.g. BH evolution, AGN feedback over cosmic time
- Steeper sources tend to lie at higher redshifts: the α -z correlation
- Known and used for decades to search for high-z objects
 - Tielens+1979, de Breuck+2000



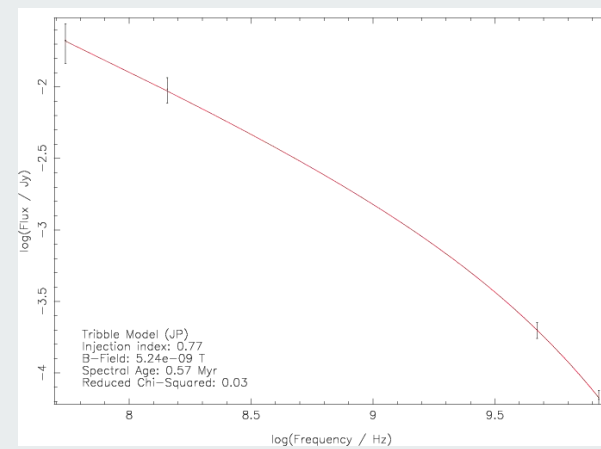
de Breuck+2000



High-z radio galaxies and the α -z correlation

- Origin has been discussed for long
 1. **Observational** effects: Malmquist bias (e.g. Blundell+1999)
 2. **Environmental** effects: environmental density (e.g. Klamer+2006)
 3. **Inverse Compton** losses (e.g. Ghisellini+2014, Morabito&Harwood 2018)

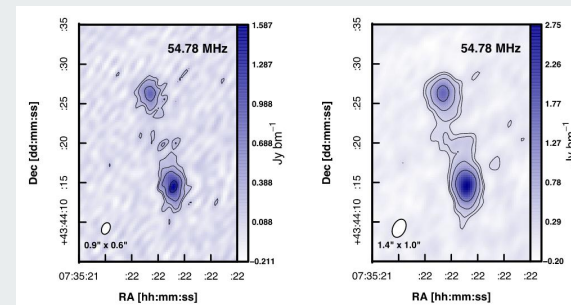
Spectral ageing



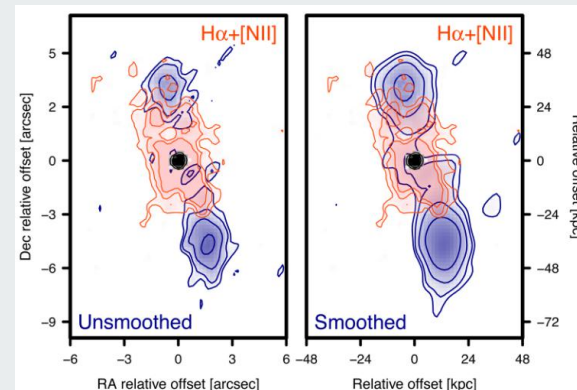
- Model energy losses as inverse Compton + Synchrotron losses
- **B**roadband **R**adio **A**stronomy **T**ool**S** (BRATS, Harwood+2013, Harwood+2015)
- Probing low (rest-frame) frequencies is important: **injection index**
- ILT uniquely probes the **low rest-frame frequencies** at **high spatial resolution**

4C 43.15

- Powerful radio galaxy at $z=2.429$ (8.3 kpc / arcsec) , $L_{143\text{ MHz}} = 2 \times 10^{29} \text{ W Hz}^{-1}$
- Selected as ultra-steep spectrum source: $\alpha = -1.1$ (325 - 1400 MHz)
- Follow up to Morabito+2016
 - Add 120-166 MHz data and study spatial distribution of spectra
 - Using LOFAR 55 MHz and 143 MHz + VLA 4.7 GHz and 8.1 GHz

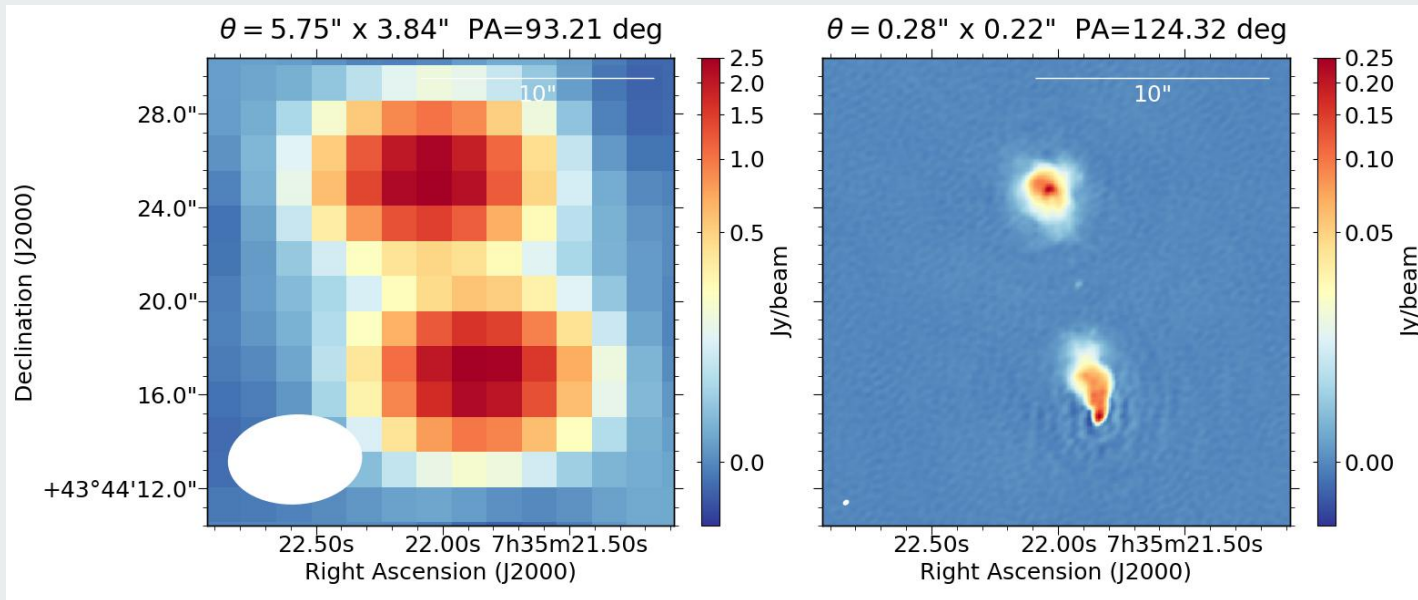


Morabito+2016



Motohara+2000, Morabito+2016

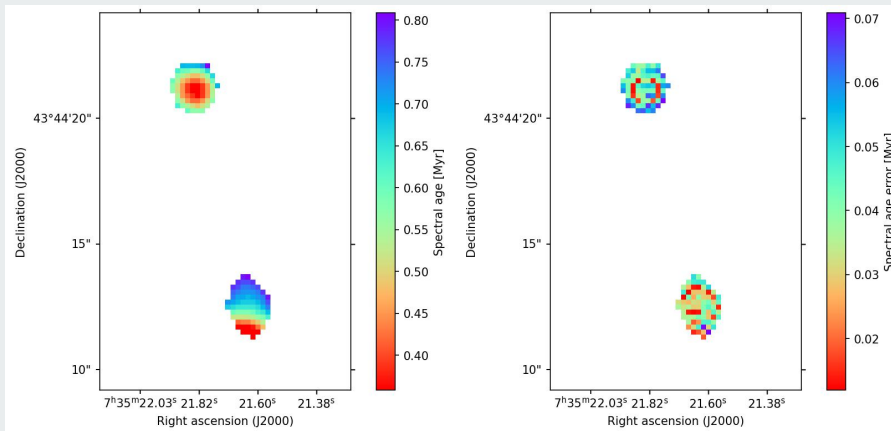
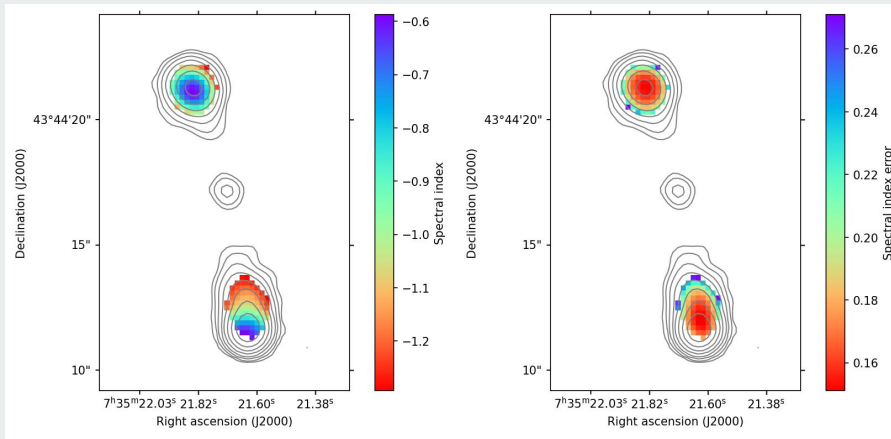
Results



Results

- Strong magnetic field, but close to CMB
 - $B = 5.2$ nT and $B_{\text{CMB}} = 3.7$ nT
- Spectral index (top) and age (bottom) distribution at $0.9''$ resolution

	North	South
Injection index	-0.8	-0.7
Spectral age [Myr]	0.81 +/- 0.06	0.90 +/- 0.6
Advance speed [c]	0.14	0.17





Conclusions

- Spatially resolved spectral index/age maps at low frequencies and high redshift
- Injection indices and advance speeds consistent with literature values
- 4C 43.15 supports the idea of increased inverse Compton losses as a driver behind the α -z correlation
 - Powerful sources are rare
 - Strong B-field close to $B / B_{\text{CMB}} = 1.5$

Thank you! Questions?